Supplementary Materials for

Sequencing seismograms: A panoptic view of scattering in the core-mantle boundary region

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Materials and Methods

Sdiff Waveforms

Broadband waveform data and synthetics were downloaded from the IRIS Data Management Center (DMC) and Instaseis databases hosted by Syngine (*36*), respectively. We used Mw > 6.5 earthquakes recorded by broadband seismic stations operating during 1990-2018. Instrument response was removed from the raw data. Tangential components of the displacement seismograms were bandpass filtered between 15 to 100 s using a Butterworth filter and windowed around the Sdiff phase (30s before and 100s after the Sdiff phase), as predicted by the 1D Earth model, PREM (*11*). Manual inspection of the selected Sdiff window was done to only include waveforms with clear Sdiff arrivals. To avoid any other bodywave traffic in the analysis window, we selected only events with hypocenter depth greater than 200 km. We then deconvolved synthetics from data using both time- and frequency-domain deconvolution, which removed radiation pattern effects and highlighted signals not predicted by PREM. We also compared our waveforms to those predicted by other 1D regional models (*28, 37*) for the Pacific basin (Fig. S2).

The Sequencer

The Sequencer is an algorithm designed to identify the leading trend in a dataset. It uses a graph-based approach to optimally order objects in a set such that the total similarity between all adjacent objects is maximized. The algorithm performs this search over a range of scales and using multiple metrics (the Euclidean Distance, the Kullback-Leibler Divergence, the Monge-Wasserstein or Earth Mover Distance, and the Energy Distance). For each of them, it estimates the level at which a continuous trend is present in the dataset by using the elongation of the corresponding distance matrix minimum spanning tree. It then aggregates information from all relevant scales and metrics by constructing an elongation-weighted distance matrix. The minimum spanning tree of this combined distance matrix provides an ordering of the objects in the dataset that corresponds to the final sequence. This approach is designed to define a view of the data, through the combination of different metrics and scales, that leads to the most elongated manifold. In contrast to the popular dimensionality reduction algorithm t-SNE (*38*, and Fig. S12), the Sequencer algorithm is fully deterministic and, in addition, it is capable of identifying scales or subsets of pixels that contribute to the signature of an underlying trend. The source code is available at http://github.com/dalya/Sequencer and an online version of the algorithm can be used at http://sequencer.org.

For our analysis, the trends were found using the Earth Mover Distance (*39*), and computed on two scales: the full waveform and 1/3 subsets. To emphasize the Sdiff postcursors and suppress apparent waveform differences due to travel-time variations, we aligned the deconvolved waveforms by cross-correlation and removed the mean trace from each waveform before sequencing. Postcursors sorted by the Sequencer become most apparent when we focus on the records after the main Sdiff positive pulse and apply histogram equalization (Fig. S13) to the records prior to inputting them into the Sequencer. The pixel-based histogram equalization here is analogous to applying automatic gain control often used in seismology to achieve optimized but nonlinear normalization. Once sequenced, we applied a running median filter to compute residual waveforms and discard as noisy those waveforms that have root-mean-square amplitudes larger than 1-sigma (Fig. S14).

We assessed Sequencer sensitivity to noise by deploying it on synthetic waveforms constructed from a real, randomly-selected waveform from our dataset contaminated by arbitrary

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but realistic noise (e.g., uncorrelated white noise drawn from a normal distribution with standard deviation of the real waveform). We found that when the noise is bandpass filtered between 15 to 100 s (like our data) and possesses a dominant period similar to the duration of the signal being analyzed, the Sequencer detects trends even from noise (Fig. S15). Users must be aware of this fact and cautious when assessing robustness to noise in order to avoid identification of spurious trends.

Significance of Postcursor Waveforms

The order determined by the Sequencer reveals trends in waveform characteristics that allowed us to identify in ~40% of waveforms coherent post-Sdiff arrivals, which we call "postcursors" (Fig. 1B, red box). To quantify the characteristics of these postcursors and how they differ from the other 60% of waveforms, we measured the delay-time and amplitude of the largest post-Sdiff arrivals, relative to the main Sdiff phase. We found that the postcursor waveforms have distinct time-amplitude relationships, including, in the vicinity of Hawaii and Marquesas, significant anti-correlation between the two (Fig. 3), which is not seen in non-postcursor waveforms. Additionally, Fig. 4B shows that the 75th percentile of amplitudes on non-postcursor paths is smaller than the 25th percentile of amplitudes on postcursor paths.

It is possible that deconvolution side-lobes contribute to waveform amplitudes after Sdiff arrival. To quantify this contribution, and rule out deconvolution as the source of the postcursor signals, we carried out transdimensional Hierarchical Bayesian deconvolution (THBD: *40*), which enables amplitude uncertainty to be estimated. We computed ensemble solutions for 4 different Sdiff waveforms, chosen to be most highly correlated with the average postcursor waveform in Hawaii (Fig. S9A) and Marquesas (Fig. S9B and S14), and average of all non-postcursor waveforms (Fig. S9D). The fourth Sdiff waveform is randomly chosen from those showing prominent postcursors near Marquesas (Fig. S9C). We used one million iterations of the THBD algorithm to deconvolve the PREM synthetic from each waveform, discarding the first half as burn-in, and saving every 100th sample to the ensemble. The posterior distributions we obtained demonstrate that postcursor amplitude cannot be an artifact due to deconvolution.

Geographic Patterns of Postcursors

Delay time and amplitude ratio measurements were made to observed postcursors with respect to the main Sdiff phases. These measurements were then assigned to the midpoint of the two D'' (the bottom 150 km of the mantle: *11*) piercing points for a given path (green, Fig. S16). For detected postcursors from the entire northern Pacific in particular, only the last 2500 waveforms (red box; Fig. 1B) in the sequence exhibit a postcursor that can be unequivocally distinguished from noise. To visualize the geographic patterns associated with delay-time and amplitude measurements and to compute bootstrap errors, we grouped postcursor measurements whose midpoints fall within a radius of 100 km of a chosen geographic location. To compute fraction of postcursors at a location, we counted waveforms whose paths travel in D'' and fall within 100 km of that location.

Postcursor measurements were also grouped in 5° bins across the Pacific basin based on the nearest point along their diffracting paths. For each geographical bin, the delay time measurements were binned in 2 s delay time windows, and the mean and its standard error were calculated (e.g., Fig. 4A). Using least squares minimization, we fit a line to the postcursor delay time and log-amplitude measurements averaged within each bin, and computed the corresponding slope and level of confidence (Fig. 3). To avoid overfitting, geographical and time bins containing three or fewer waveforms were excluded from the analysis.

Waveform Simulations

We tested various scenarios of the lowermost mantle by simulating waveforms using the three-dimensional spectral element method (*41*). The moment tensor solution for each event used in the modeling was obtained from the GlobalCMT catalog (*42*). We explored two types of models: ULVZs and sides of LLVSPs. We systematically varied the height and radius of, and wavespeed reduction in, a cylindrical ULVZ model (Fig. S6). We simulated head-on and edge-on scattering from the sharp sides of the LLSVPs, assuming a 5% Vs contrast (*9*) and large wavelength sinusoidal boundary undulations. For both cases, we used cosine tapers to implement smooth velocity variations across boundaries. Other models used for the forward modeling included PREM, SEMUCB-WM1, Hawaii ULVZ (*4*), a hypothetical Gaussian plume model 600 km in height, and a sharpened SEMUCB-WM1. To sharpen SEMUCB-WM1, we linearly interpolated shear velocity perturbations in the bottom 600km of the mantle and mapped them to an arctangent function of maximum and minimum value of +5% and −5%, respectively. The height of 600 km above the CMB is approximately the maximum height at which the average Vs of LLSVP regions substantially differs from non-LLSVP regions (*18, 43*).

Differences in Effects of Fast and Slow Anomalies

High velocity anomalies can also scatter seismic waves, and may produce some of the weak-amplitude postcursors (pervasive, Fig. 1C) we observe, since these could be attributable to scattering from a compact anomaly. The normalized Rayleigh scattering coefficient for SH waves is unity (*44*), since our datasets are predominantly in the forward scattering regime. Therefore, an appropriately located high-velocity anomaly could produce postcursors similar to those from a low velocity anomaly (Fig. 4B).

However, generating the large amplitude postcursors, particularly beneath Hawaii and Marquesas, requires low-velocity compact anomalies based on two main reasons. First, low velocity anomalies can amplify the postcursors by focusing. Previous work has shown that focusing of seismic energy by a slow anomaly may be necessary to explain the large postcursor amplitude and their moveout (*4*). Second, waves transmitted through the anomalies and delayed by them play a significant role producing such large amplitude postcursors. Based on our waveform simulations, a low velocity cylindrical structure ~1000 km across and with a 20% Vs reduction can produce postcursors with an amplitude ratio greater than 1 with 25-30 sec time day (Fig. 4A). This is because what started out as the original Sdiff wavefront is delayed to produce postcursors, while what we interpret as the main Sdiff arriving before the postcursors is actually a relatively weaker phase produced by wavefront healing. On the other hand, a similar structure but with a positive velocity anomaly (20% Vs increase) is only capable to produce much smaller amplitude (~0.25) postcursors (Fig. 4A). In this case, what starts out as the original Sdiff wavefront remains the first-arriving Sdiff energy, and the relatively weaker delayed postcursor energy is produced by waves diffracting around the fast anomaly.
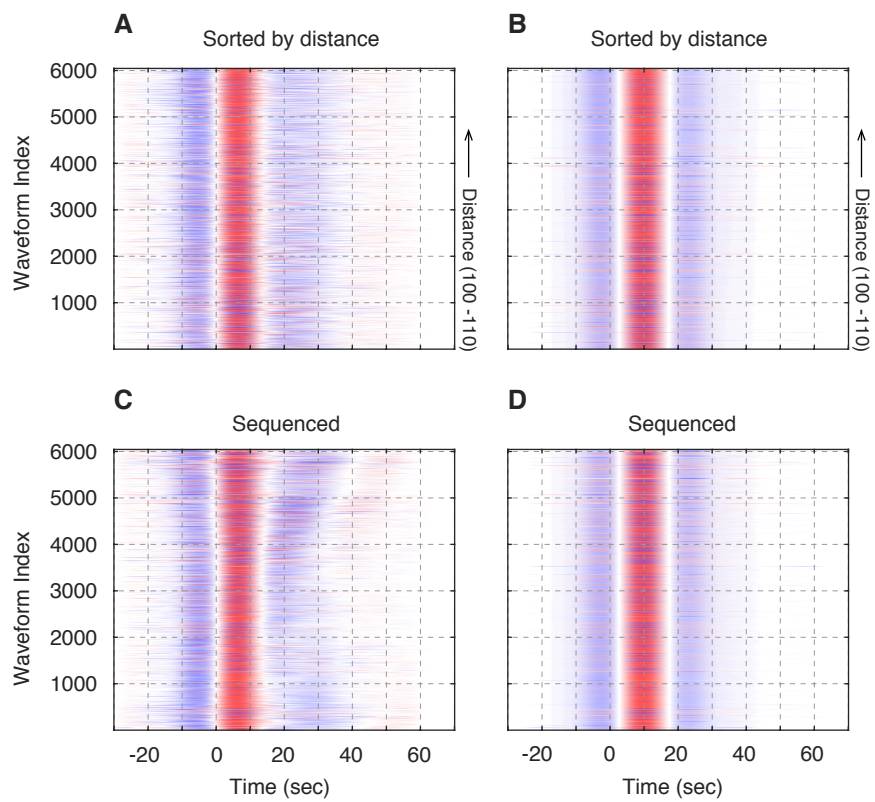
A Sorted by distance

B Sorted by distance

C Sequenced

D Sequenced

Fig. S1. Raw tangential seismograms (A) and synthetic waveforms computed for PREM (B) sorted by distance. (C-D) Same as A and B but sorted by the Sequencer.
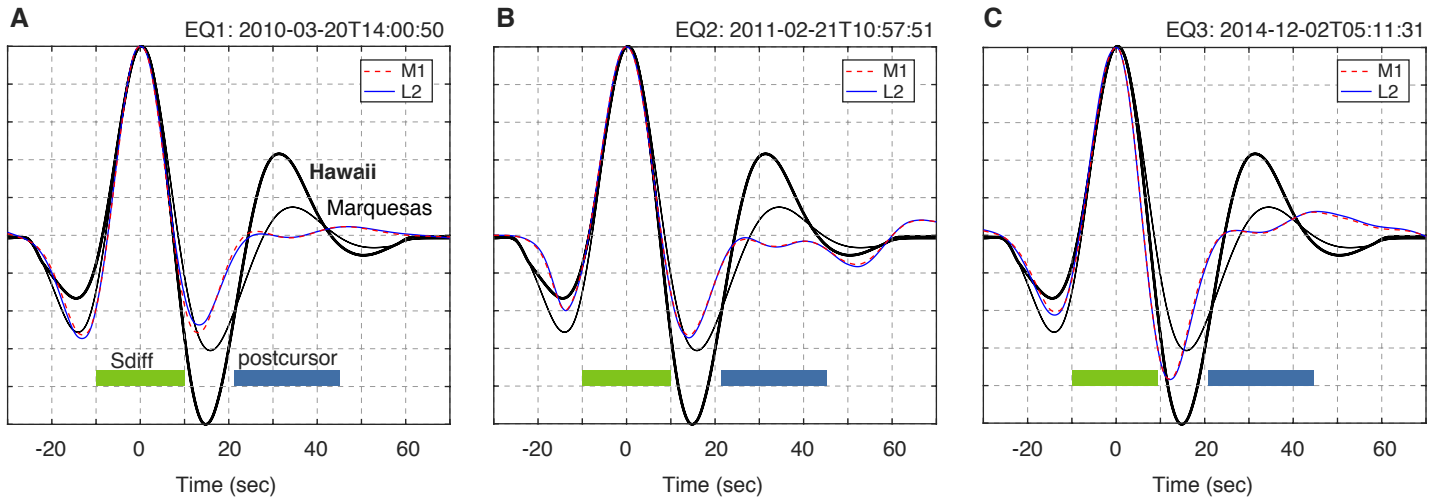
Fig. S2. (A-C) Comparison of average postcursor waveforms from Hawaii and Marquesas against average synthetic waveforms computed using two regional 1D models for the Pacfic basin (*28*, *37*). Three events used in the waveform simulation are shown in Fig. S8. At both hotspots, strong postcursors (indicated by blue horizontal bar) are apparent 30-40s after the main Sdiff arrival (green) for all three events, which are not predicted by the 1D models.

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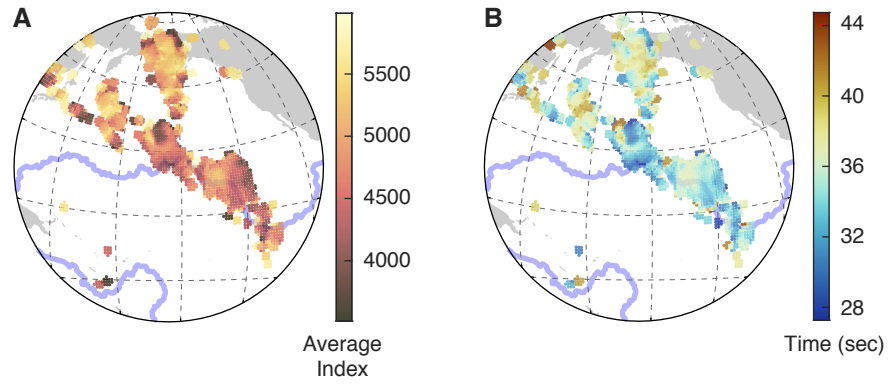
Fig. S3. Stack of average sequencing index and postcursor delay time relative to main Sdiff arrival averaged in 1° bins. The geographic extent of the Pacific LLSVP (*18*) is shown (light blue contour).
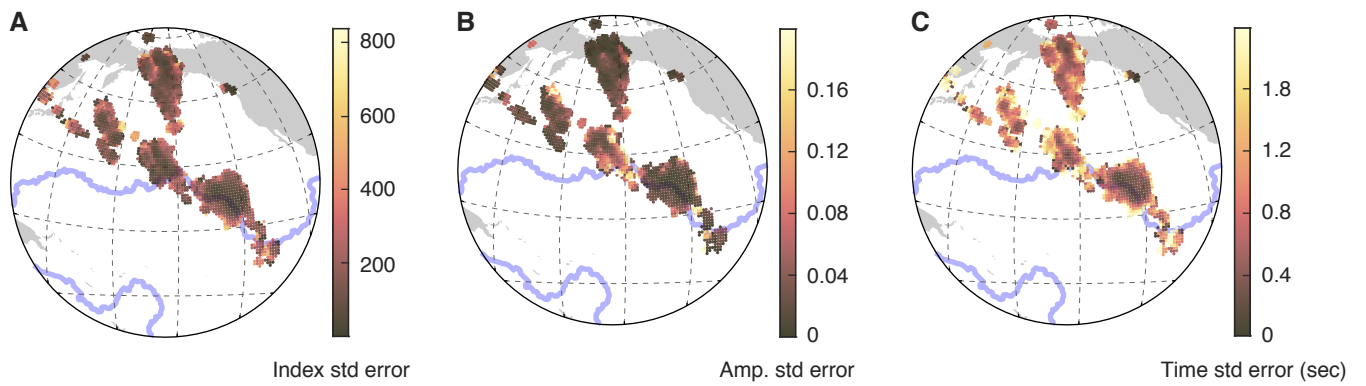
Fig. S4. Standard errors of the postcursor (A) sequencing index, (B) amplitude and (C) delay time relative to the main Sdiff arrival. The geographic extent of the Pacific LLSVP (*18*) is shown (light blue contour).
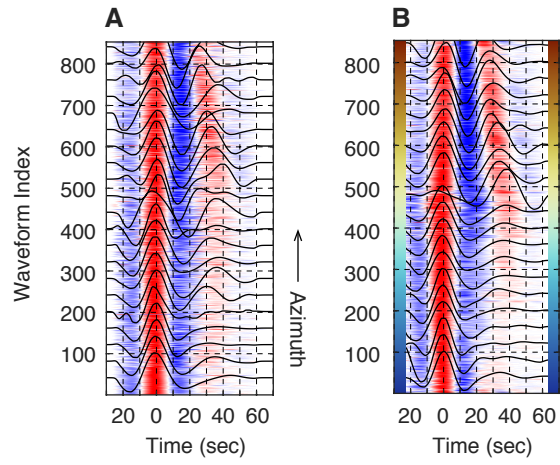
Fig. S5. Deconvolved Sdiff waveforms with turning points within 20° of Hawaii sorted by (A) azimuth and (B) the Sequencer.
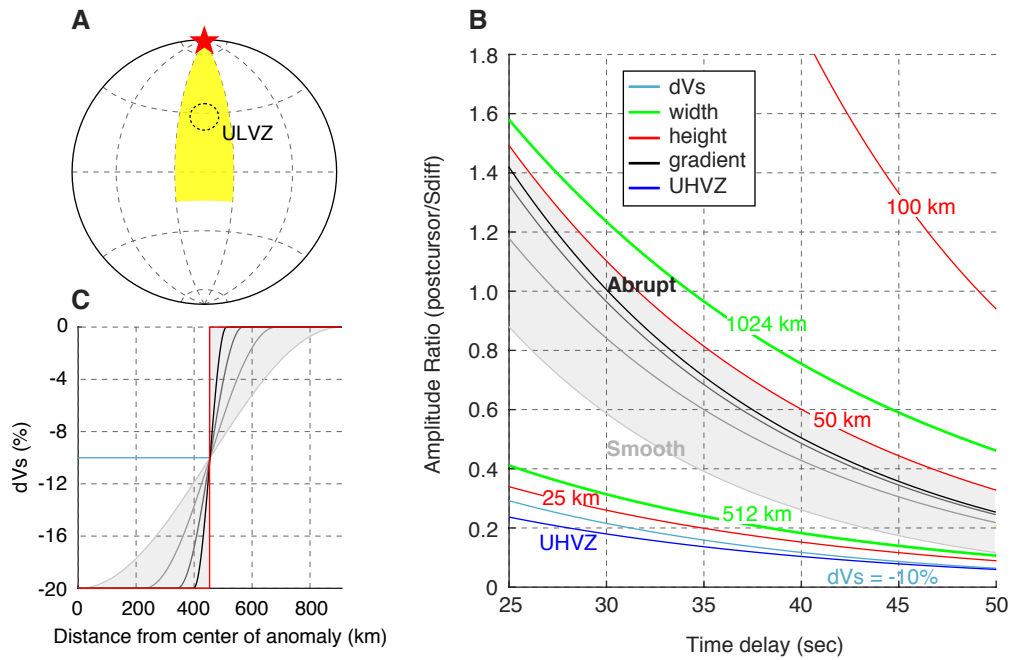
Fig. S6. Effect of width, height, velocity reduction, and velocity gradient on postcusor delay time vs. amplitude trends for a cylindrical velocity anomaly. (A) Illumination geometry. (B) Postcursor delay time and amplitude measured on deconvolved SPECFEM synthetics, with lines labeled according to the parameter being varied with respect to a reference model with: dVs = -20%; width = 910 km; height = 50 km; gradient = none. The ultrahigh-velocity zone (UHVZ) is identical to the reference model, except dVs = +20%. (C) As the gradient parameter increases from none (red), the velocity change across the edge of the cylindrical anomaly becomes smoother.
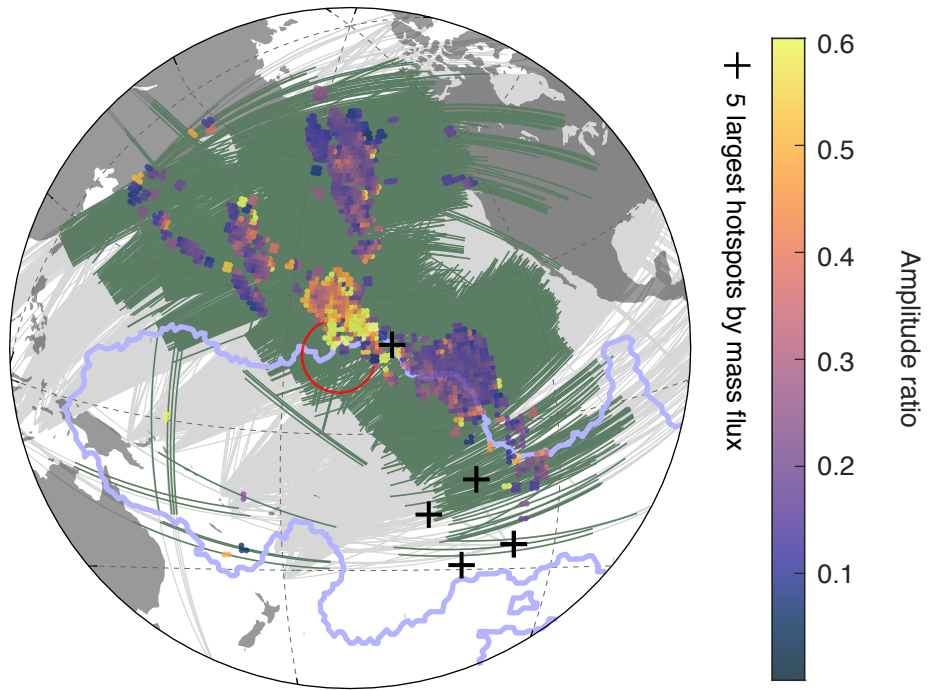
Fig. S7. Sdiff paths in D'' beneath the Pacific basin (dark green). The mega-ULVZ model (*4*) is outlined in red. All other symbols are same as in Fig. 1D.
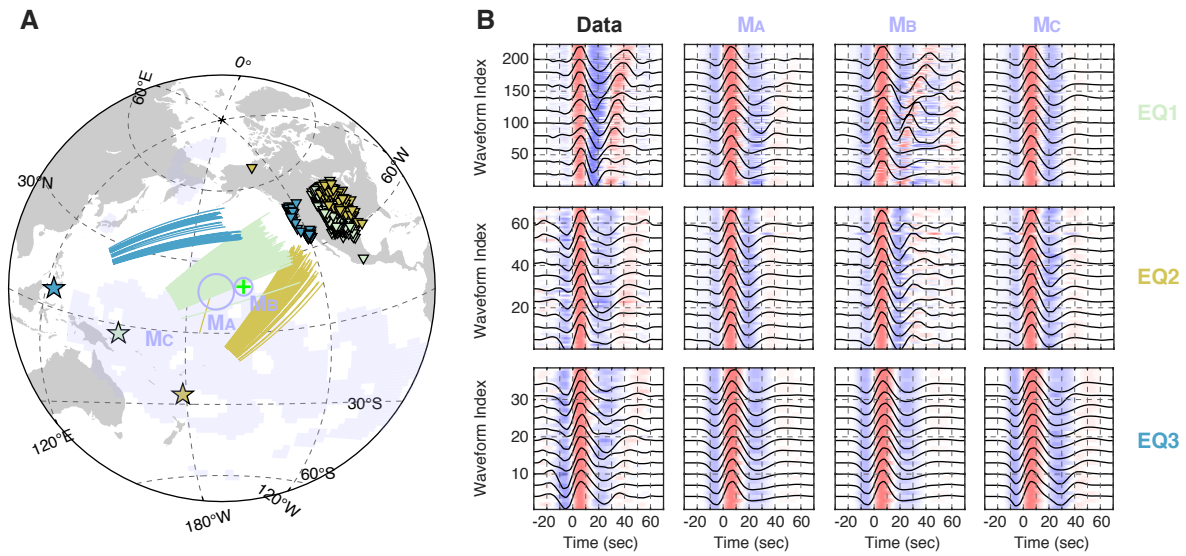
Fig. S8. Comparison of observed waveforms from three earthquakes with their corresponding synthetics computed for three different lowermost mantle structures. M_A is the mega-ULVZ model from (*4*), M_B is a 600 km tall, hypothetical gaussian plume centered beneath the Hawaiian hotspot (green cross), and M_C denotes a sharpened SEMUCB-WM1 (*30*) in the bottom 600 km of the mantle. See (*10*) for details. Earthquakes, stations, and paths are displayed on the map. (B) Sequenced deconvolved waveforms from both data and models aligned on the main Sdiff arrival.
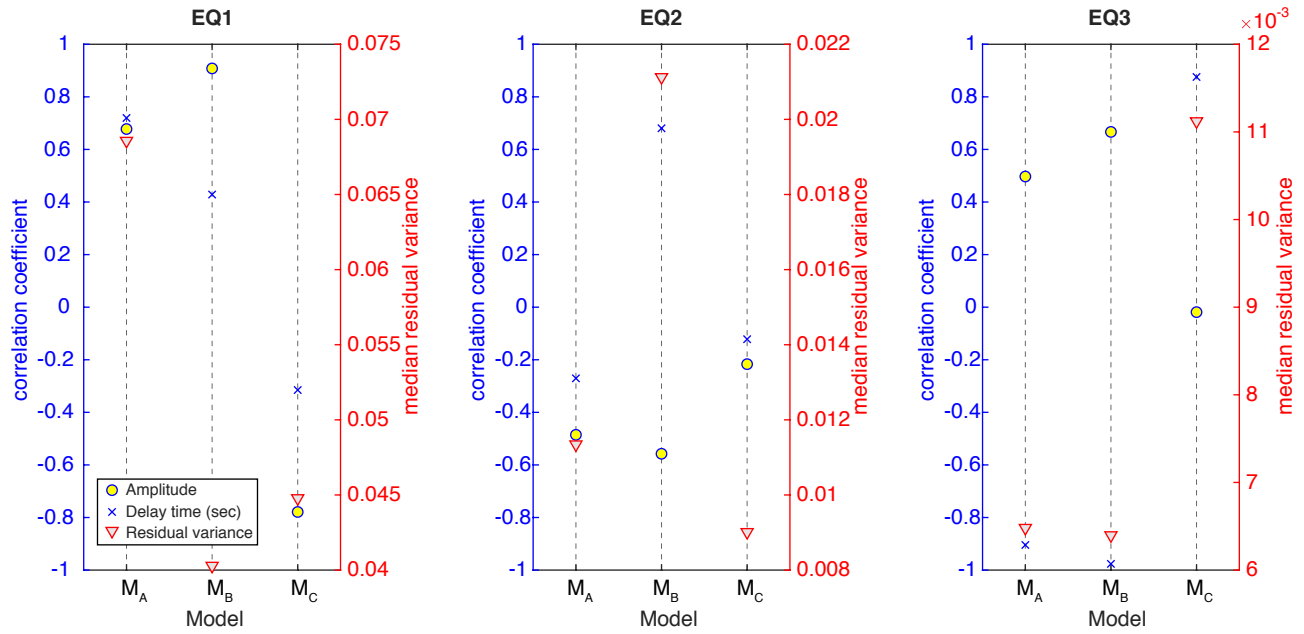
Fig. S9. Model fits for data vs. synthetic waveforms in Fig. S8. For each earthquake, we plot the correlation coefficients and residual variance (triangles) between data and synthetics comptued for each model. Correlation coefficients are computed separately for postcursor amplitude (circles) and delay time (crosses).
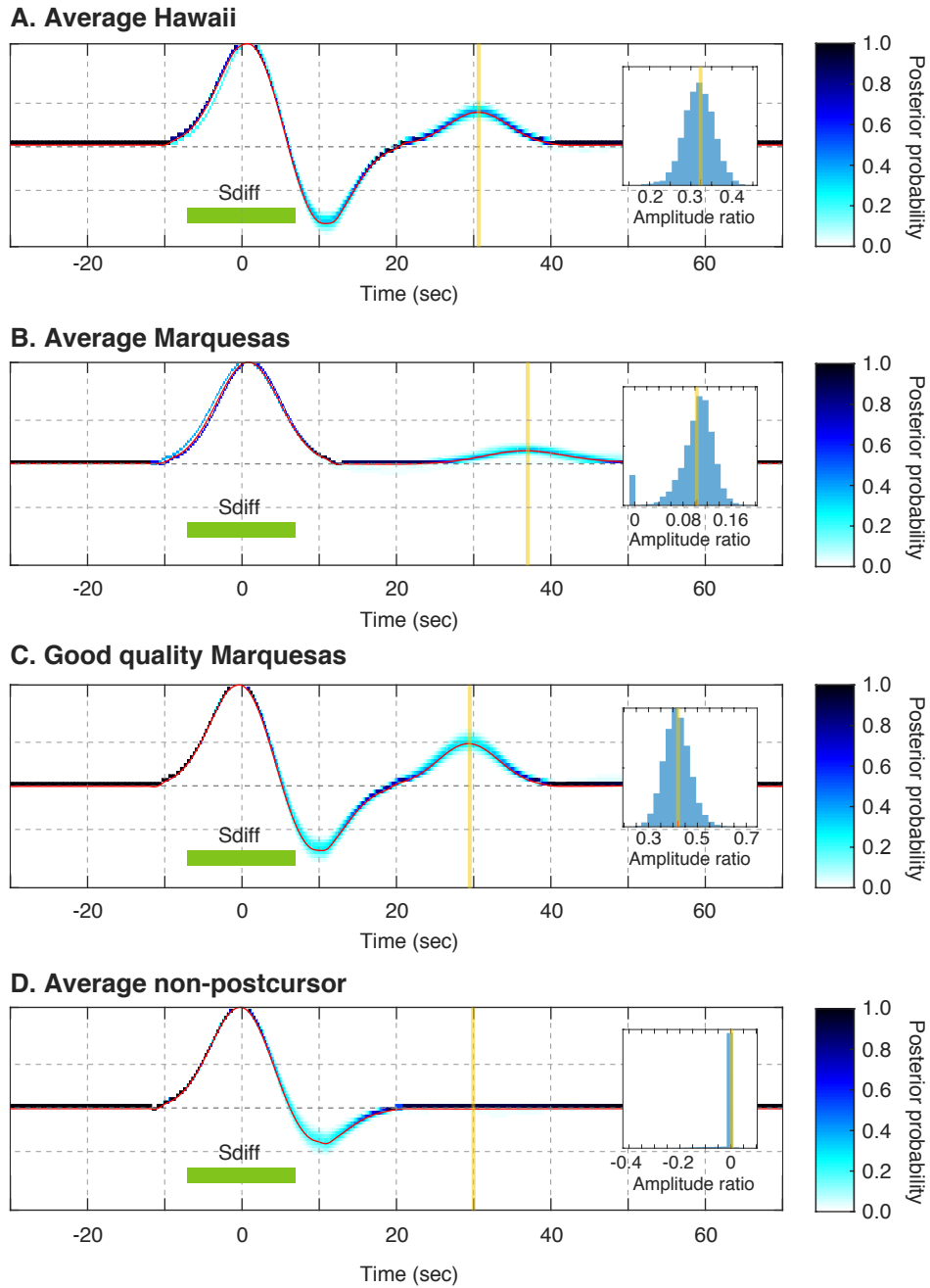
A. Average Hawaii

B. Average Marquesas

C. Good quality Marquesas

D. Average non-postcursor

Fig. S10. Quantitative assessment of observed postcursors amplitudes. Ensemble solutions for deconvolved Sdiff waveforms and their associated uncertainties in (A) Hawaii, (B-C) Marquesas, and (D) those that do not show postcursors. One million iterations were carried out for each waveform using a transdimensional hierarchical Bayesian deconvolution (*40*). Yellow vertical line indicates the time at which the post-Sdiff amplitude is observed, and the inset shows the corresponding conditional probability density for amplitude.
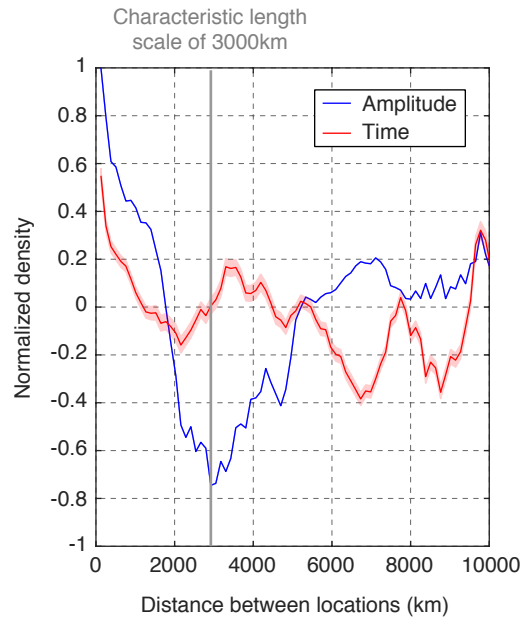
Fig. S11. Two-point correlation functions for amplitude (blue) and delay time (red) of postcursors shown in Fig. 1B. The correlation function for postcursor amplitude (blue line) is smallest at ~3000 km, corresponding to the characteristic length scale of the spatial pattern of postcursor amplitude variation across the Pacific basin.
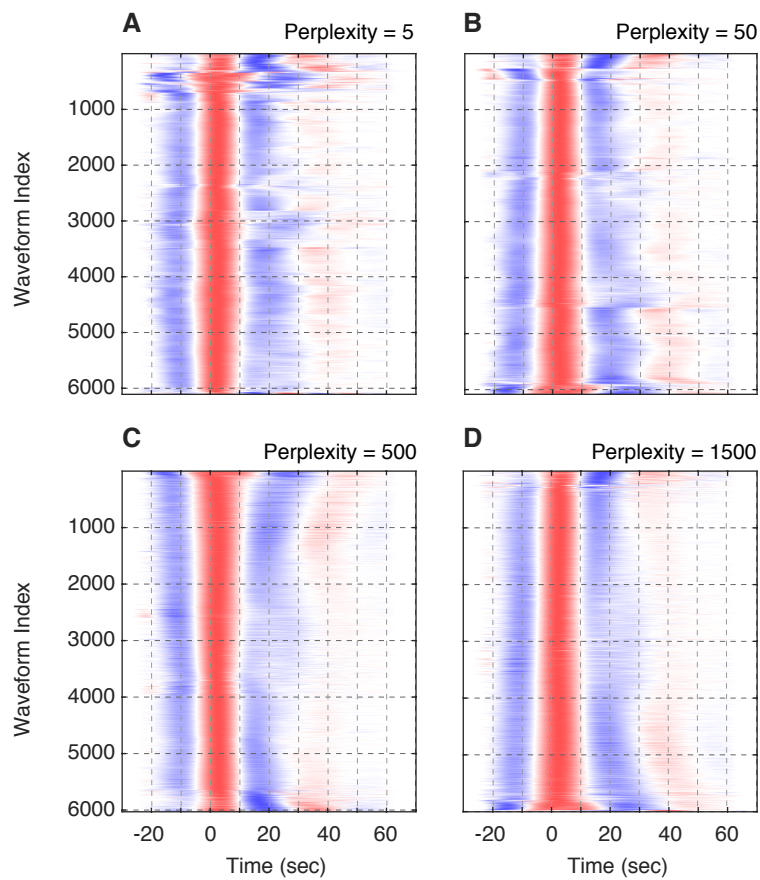
Fig. S12. Waveforms in Fig. 1B sorted by the t-SNE algorithm (*38*). Note that using the Sequencer is superior in detecting a clear trend of the postcursors (Fig. 1B). Although postcursors get more apparent with larger values of perplexity (A-D), the exact indices cannot be reproduced due to the stochasticity in t-SNE.
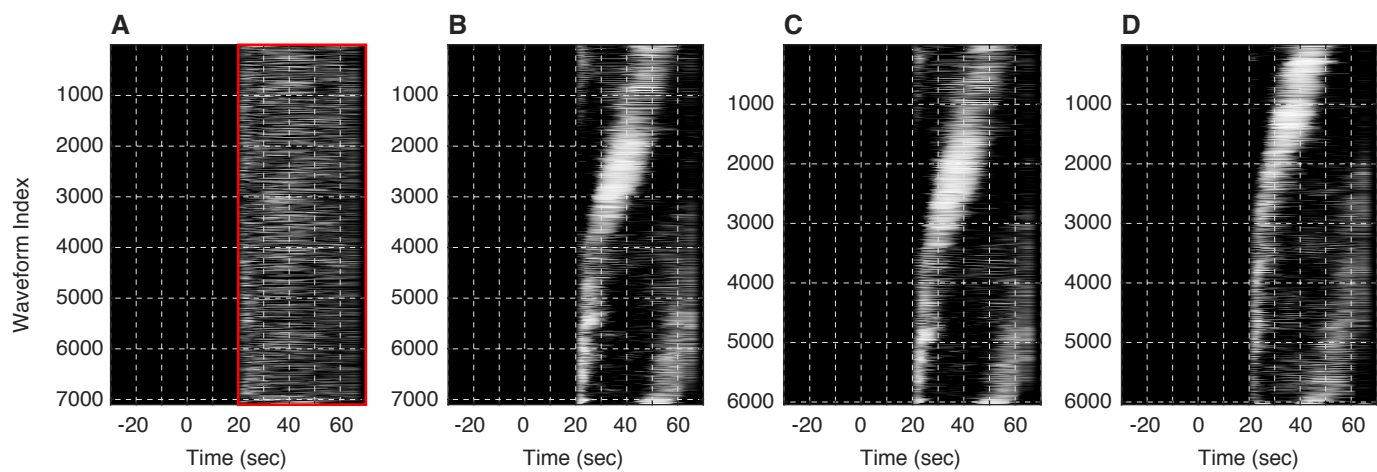
Fig. S13. Histogram equalized waveforms. (A) Historgram-equalized data after the main Sdiff positive pulse (red box) are used as the initial input for the Sequencer. (B) Same as A but sequenced. (C) Same as B but with outliers discarded according to procedure in Fig. S14. (D) Final sequence determining the order of waveforms shown in Fig. 1B.
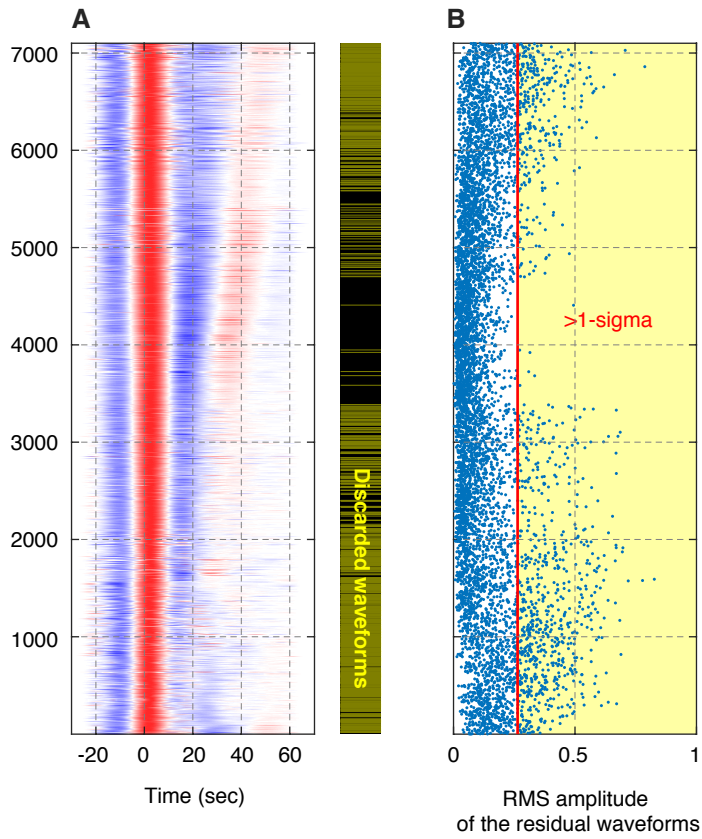
Fig. S14. Discarding outliers. (A) Sequenced Sdiff waveforms prior to discarding noisy data. Horizontal yellow bars indicate waveforms that are discarded. (B) RMS amplitudes of the residual waveforms with respect to a moving median of the sequenced waveforms. Red vertical line is 1-sigma RMS amplitude and waveforms to its right are discarded.
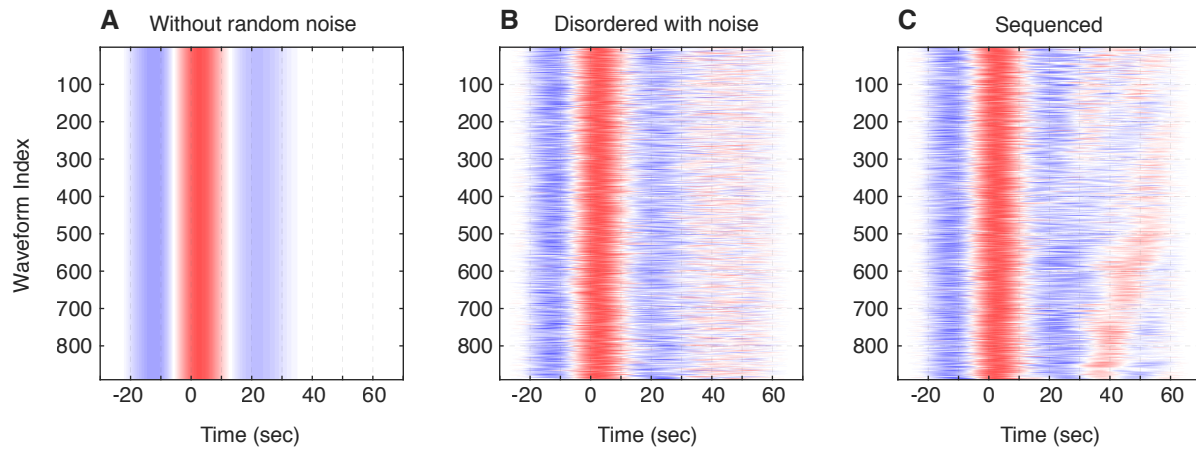
Fig. S15. (A) Randomly-selected waveform from Sdiff dataset. (B) Uncorrelated white noise added to A. (C) Same as B but sorted by the Sequencer.
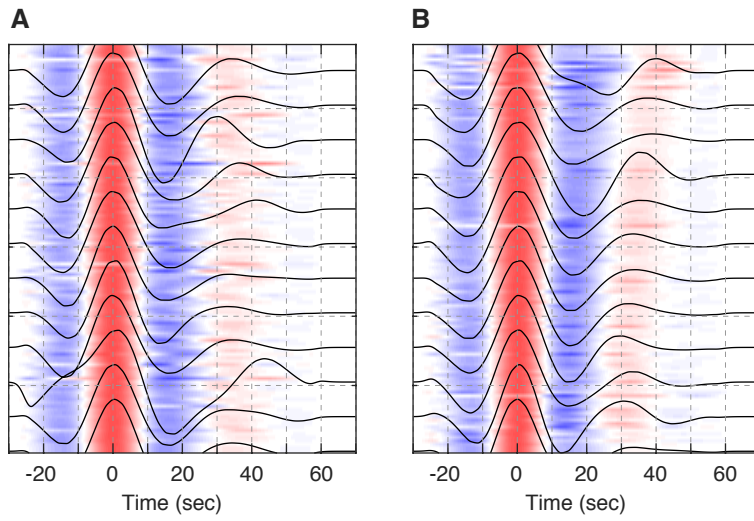
Fig. S16. Postcursors in the Marquesas region. Deconvolved Sdiff waveforms with turning points within 5° of Marquesas sorted by (A) distance and (B) the Sequencer.

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